

## Übungen zur Vorlesung Relativitätstheorie und Kosmologie II: Problem Sheet 4

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- 1 Null Geodesics in Schwarzschild Let  $\gamma$  be an affinely parameterized null geodesic for the Schwarzschild metric with  $J \neq 0$  lying in the equatorial plane  $\theta = \pi/2$ .

Q1. Show that along  $\gamma$ , both for  $u > 1/2$  and for  $u < 1/2$ , we have

$$p^2 = 2u^3 - u^2 + \alpha^2, \quad (1)$$

where  $\alpha$  is a constant and

$$u = \frac{m}{r}, \quad p = \frac{du}{d\phi}.$$

Q2. For the case  $\alpha = 0$  show that we obtain an autonomous two-dimensional system

$$\frac{du}{d\phi} = p, \quad \frac{dp}{d\phi} = 3u^2 - u.$$

Find and classify the critical points. Sketch the trajectories in the  $(u, p)$  phase-plane.

Q3. For the case  $\alpha = 0$  and  $u > 1/2$  show that the geodesic has an equation of the form

$$r = 2m \cos^2\left(\frac{\varphi - \varphi_0}{2}\right), \quad (2)$$

with  $t = t(\varphi)$  that you should determine.

- 2 Timelike geodesics in Schwarzschild Let  $\gamma$  be a timelike geodesic for the Schwarzschild metric parameterized by proper time and lying in the equatorial plane  $\theta = \pi/2$ .

Q1. Show that

$$\frac{E^2 - \dot{r}^2}{1 - \frac{2m}{r}} - \frac{J^2}{r^2} = 1.$$

Q2. Deduce that if  $E = 1$  and  $J = 4m$  then

$$\frac{\sqrt{r} - 2\sqrt{m}}{\sqrt{r} + 2\sqrt{m}} = Ae^{\epsilon\varphi/\sqrt{2}},$$

where  $\epsilon = \pm 1$  and  $A$  is a constant. Describe the orbit that starts at  $\varphi = 0$  in each of the cases (i)  $A = 0$ , (ii)  $A = 1$ ,  $\epsilon = -1$ , (iii)  $r(0) = 3m$ ,  $\epsilon = -1$ .